# Radicals Validation Subgroup

# Ross Salawitch Jet Propulsion Laboratory, California Institute of Technology

#### 6 Presentations

NO<sub>2</sub>: OMI column, ground based validation, Table Mtn, Ca. – S. Sander OMI column, ground based validation, Tri-Cities, Wa. – E. Spinei HIRDLS profiles vs ACE profiles – C. Randall

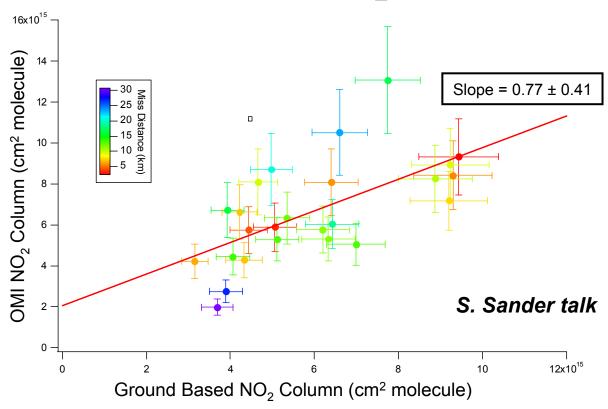
BrO: OMI global column observations – T. Kurosu MLS profiles, balloon-based validation – L. Kovalenko

OH &  $HO_2$ :

MLS Profiles, balloon-based validation – H. Pickett and T. Canty

Notes: CIO will be discussed in chlorine breakout, Wed afternoon Important "validation issues" highlighted in red

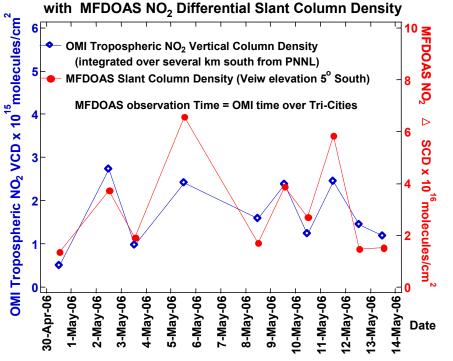
# OMI NO<sub>2</sub>



- Mountain topography and air circulation require close coincidence with OMI footprint for proper comparison
- Reasonable correlation between OMI and ground based column NO<sub>2</sub>
- OMI NO<sub>2</sub> is 20 to 30% less than ground based NO<sub>2</sub>
- Similar results seen for Tri-City, Washington comparisons ⇒ Spinei poster and for independent OMI comparisons ⇒ Gleason talk

# OMI NO<sub>2</sub>

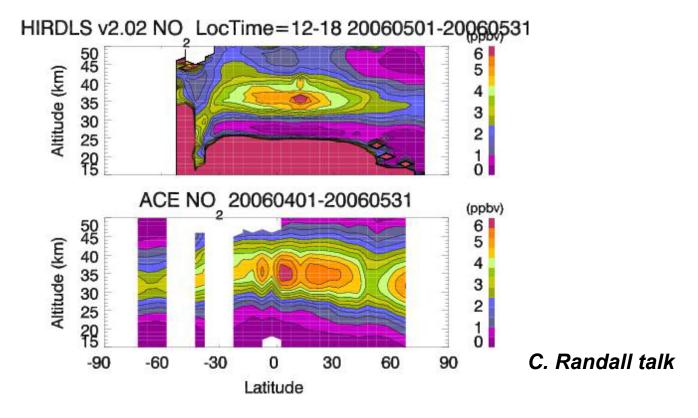
Figure 8. Comparison of OMI Tropospheric NO<sub>2</sub> Vertical Column



E. Spinei talk/poster

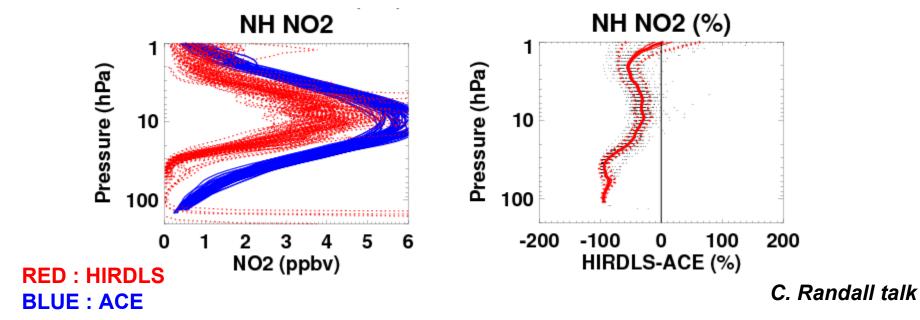
- European inter-comparison of ground based (multiple instruments) and space based NO<sub>2</sub> ⇒ occurring now!
- Inter-comparison campaign for NO<sub>2</sub> and O<sub>3</sub> at Table Mountain, Ca. including JPL FTUVS, WSU MF-DOAS, and GSFC instruments being planned for ~ mid-May to mid-June 2007
  - Other groups welcome
  - Coordination with Aura, TC<sup>4</sup> test flights, possibly SCIAMACHY desired

### HIRDLS NO<sub>2</sub> compared to ACE NO<sub>2</sub>



- NO<sub>2</sub> retrieved from HIRDLS radiances, accounting for kapton emission
- Initial comparison to ACE shows similar patterns
  - ⇒ very promising

# HIRDLS NO<sub>2</sub> compared to ACE NO<sub>2</sub>

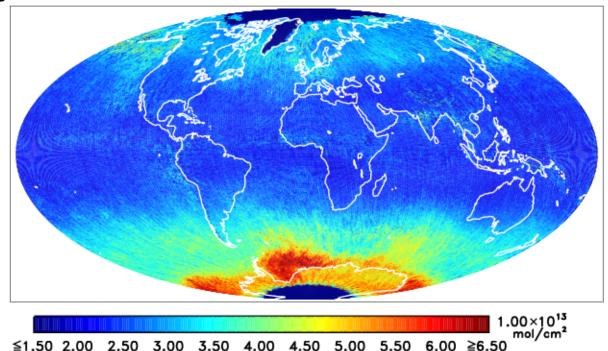


- Inter-comparisons at high latitudes only
- HIRDLS NO<sub>2</sub> lower than ACE by ~10 to 100%
- Note:
  - ACE NO<sub>2</sub> lower than HALOE NO<sub>2</sub> by 0 to 10%
  - Comparisons not yet factoring in "time of day"
     (NO<sub>2</sub> vs SZA goes in "right direction" for explaining some of the differences)

Improvements of HIRDLS retrieval, focusing on kapton correction, underway

#### Global BrO from OMI

**BrO August 2006 – cloud fraction < 20%** 

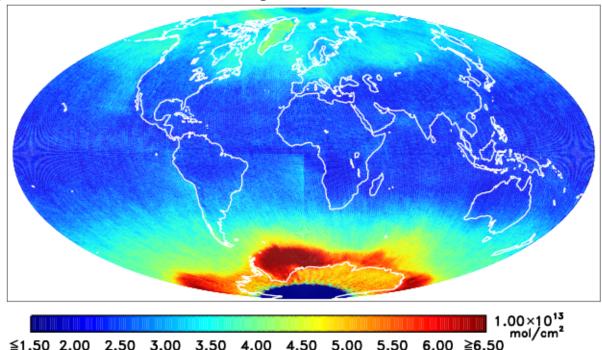


T. Kurosu talk

- Must use Version 0.9.50 or later!
- First public release: Oct 2006
- BrO columns compare favorably to GOME columns (prior years)
- Release of BrO from ice shelf and salt lakes clearly seen by OMI
- Volcanic release being studied: SO<sub>2</sub> signal must be separated

#### Global BrO from OMI

BrO August 2006 - no cloud screening



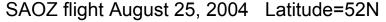
T. Kurosu talk

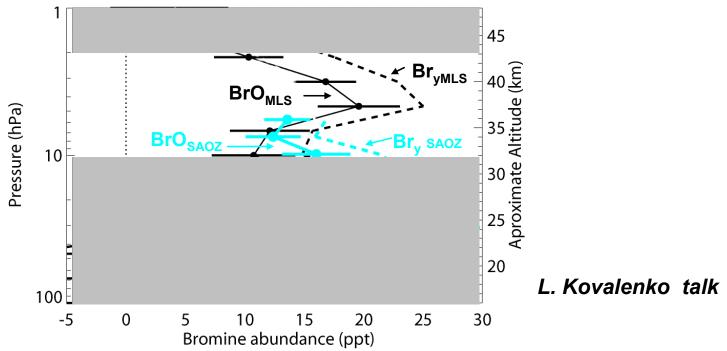
#### Future steps:

- sensitivity of BrO retrieval to O<sub>3</sub>, NO<sub>2</sub> absorptions
- understanding correlation of BrO with high albedo
- tropospheric vs stratospheric contributions to column BrO
  - ⇒ bromine budget and role of VSL bromocarbons

Empirical "OMI-based" quantification of tropospheric vs stratospheric contributions to column BrO will be attempted. Nonetheless, acquisition of aircraft BrO profiles in OMI footprint is an outstanding, as yet unachieved validation need

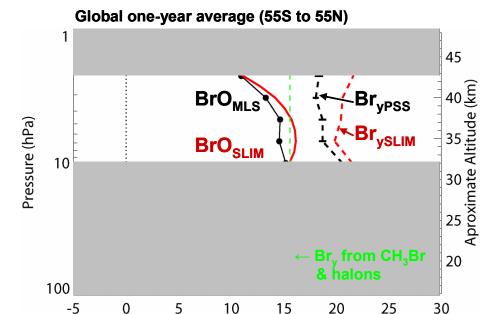
### Upper stratospheric BrO from MLS





- Three versions of MLS BrO:
  - 1.5 : not useful for scientific analysis
  - BinRad: research adjunct of 1.5, useful for scientific analysis from 2 to 10 hPa
  - 2.1: looks better than 1.5, but not as good as BinRad
- MLS (BinRad) and SAOZ (balloon) BrO and Br<sub>y</sub> agree to within respective uncertainties, 2 to 10 hPa
- MLS and DOAS (balloon) BrO also agree within uncertainties

### Upper stratospheric BrO from MLS



Bromine abundance (ppt)

L. Kovalenko talk & N. Livesey paper

First science result:

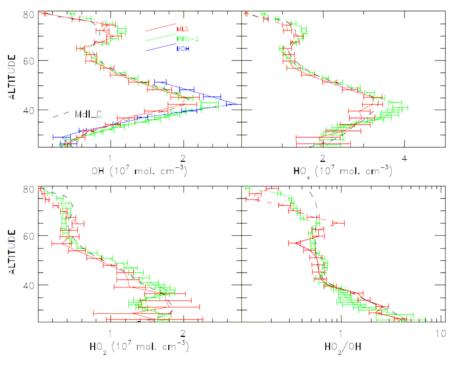
- $Br_v = 18.6 \pm 5.5 \text{ ppt}$
- VSL bromocarbon contribution to  $Br_y \Rightarrow 3.0 \pm 5.5 \text{ ppt}$

Livesey et al., GRL, accepted, 2006

#### Near Future:

- Comparison of Vers 2.1 BrO with SLIMCAT, DOAS, SAOZ, & SCIAMACHY
- Extend BrO profile to higher altitudes using model day/night differences
- Use of stratospheric BrO profiles in analysis of column BrO (i.e., synergistic analysis of MLS and OMI BrO)

### Upper stratospheric HO<sub>x</sub>: Sept 2005

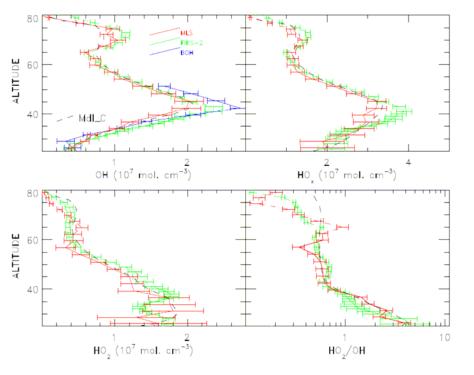


H. Pickett & T. Canty talk & T. Canty poster;

Update to:
H. Pickett et al.
& T. Canty et al.
GRL 2006 papers

- Figures show OH, HO<sub>2</sub>, HO<sub>x</sub>, and HO<sub>2</sub>/OH from:
  - MLS Vers 2.1 (red), two balloon instruments (FIRS-2 and BOH) (green/blue), and model (black)
- Vers 2.1 provides:
  - higher vertical resolution retrieval of mesospheric OH
  - smoother stratospheric retrieval HO<sub>2</sub>
- Greater discrepancy between MLS and balloon OH near 40 km during Sept 2005 than seen during Sept 2004
  - reasons under study and not yet understood, but <u>not</u> due to updated MLS retrieval version
  - Sept 2004 profiles were subject of published HO<sub>x</sub> validation paper (Pickett et al., GRL, 2006)

### Upper stratospheric HO<sub>x</sub>: Sept 2005



H. Pickett & T. Canty talk & T. Canty poster;

Update to:
H. Pickett et al.
& T. Canty et al.
GRL 2006 papers

Initial science results, based on MLS Vers 1.5 and Sept 2004 balloon data:

- No indication of previously noted "HO<sub>x</sub> dilemma" in comparison of modeled & meas HO<sub>x</sub> profiles
- Change to  ${\rm HO_x}$  kinetics resulting in "best agreement" with  ${\rm HO_x}$  does not resolve upper stratospheric "ozone deficit problem"
- Above results based on MLS Vers 1.5 profiles, Sept 2004 to June 2005 & Sept 2004 balloon data (Canty et al., GRL, 2006)

Preliminary analysis of MLS Vers 2.1 profiles for Sept 2005 yields same scientific conclusions as above (Canty et al. poster)